

KIC 003954211

Q1-17 DR25 TCE Parameters

| TCE | Run Type | KOI? | Period (Days) | Epoch (BKJD) | Depth (ppm) | Duration (Hours) | MES | SNR | R_{\star} (R_{\odot}) | T_{\star} (K) | R_p (R_{\oplus}) | S_p (S_{\oplus}) |
|--------------|----------|---------|---------------|--------------|-------------|------------------|-------|------|-----------------------------|-----------------|------------------------|------------------------|
| 003954211-01 | OBS | 8090.01 | 0.594276 | 131.837901 | 592.1 | 1.234 | 154.0 | 36.9 | 1.12 | 6321 | 5.12 | 8346.19 |

Robovetter Results

| TCE | Run Type | Disp | Score | N | S | C | E | Comments |
|--------------|----------|------|-------|---|---|---|---|----------------------------------|
| 003954211-01 | OBS | FP | 0.00 | 0 | 0 | 1 | 1 | CENT_RESOLVED_OFFSET—EPHEM_MATCH |

Notes: OBS = Observed. INJ = Injected. INV = Inverted. SCR = Scrambled.

N = Not Transit-Like. S = Stellar Eclipse. C = Centroid Offset. E = Ephemeris Match.

See http://exoplanetarchive.ipac.caltech.edu/docs/API_kepcandidate_columns.html#proj_disp_col for comment definitions.

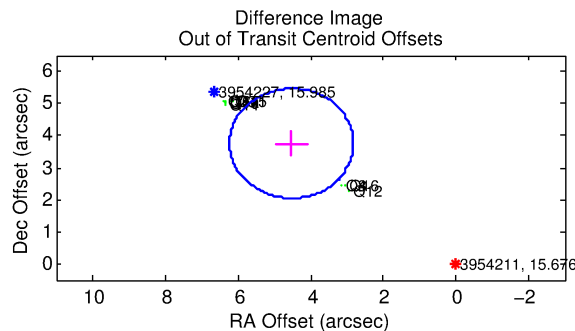
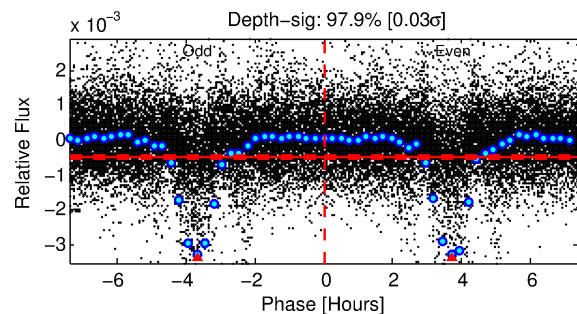
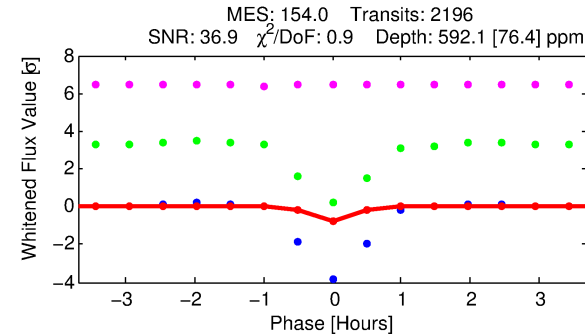
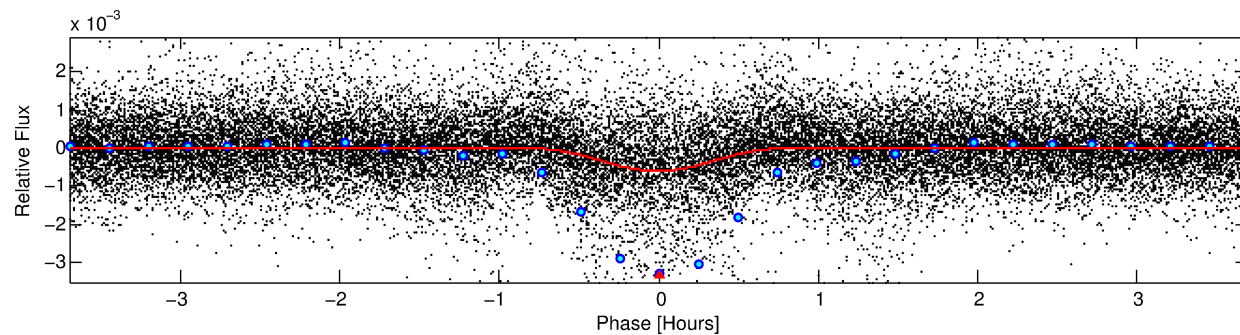
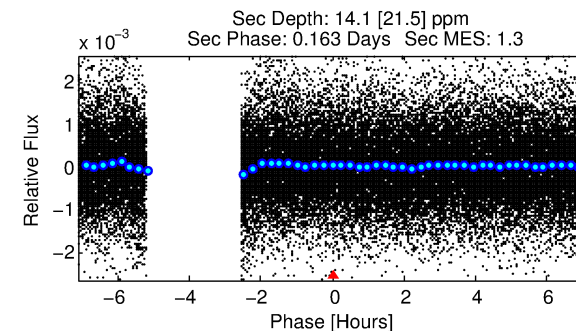
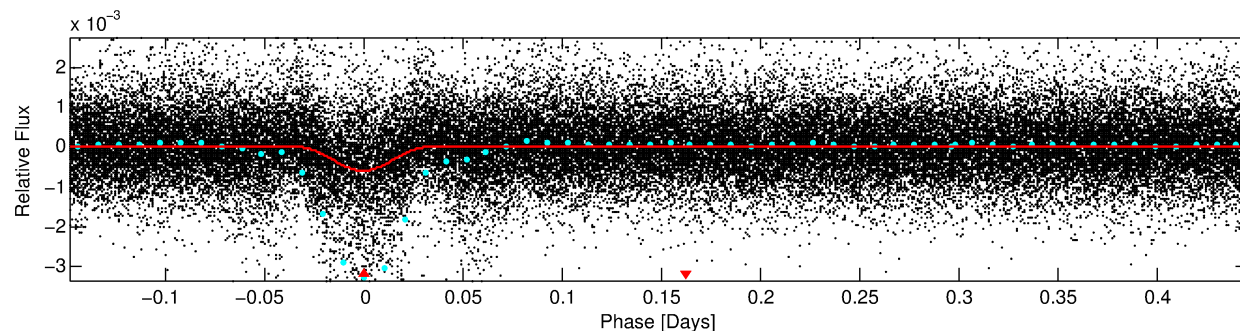
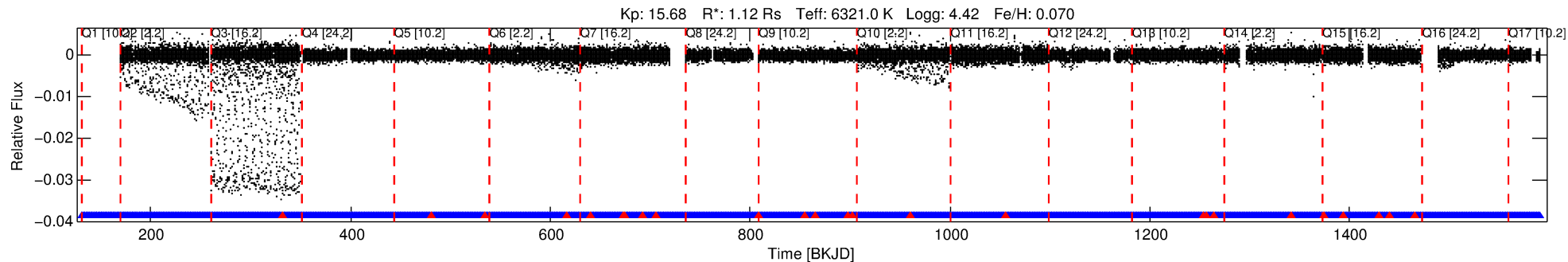
Ephemeris Match Information For 003954211-01

| TCE (1) | KIC | Parent (2) | Parent KIC | $P_1:P_2$ | Dist ($''$) | Δ Row | Δ Col | m_2 | m_1 | D_2/D_1 | Mechanism | Flag | σ_P | σ_T |
|--------------|---------|------------|------------|-----------|---------------|--------------|--------------|-------|-------|-----------|------------|------|------------|------------|
| 003954211-01 | 3954211 | 3766.01 | 3954227 | 1:1 | 8.6 | 1 | 2 | 15.98 | 15.67 | 372.87 | Direct-PRF | 0 | 0.01 | 0.23 |

Notes: $P_1:P_2$ is the period ratio. Dist is the distance in arcseconds. Δ Row and Δ Col are the number of pixels apart in row and column. m_2 and m_1 are the magnitudes of the parent and child. D_2/D_1 is the parent's transit depth divided by the child's. σ_P and σ_T are the significance of the match in period and epoch. For a match to be considered significant $\sigma_P < 5.0$ and $\sigma_T < 5.0$. Matches which have σ_P and σ_T very close to this cutoff should receive extra scrutiny, especially if the period ratio is very large.

DV One-Page Summary

KIC: 3954211 Candidate: 1 of 1 Period: 0.594 d



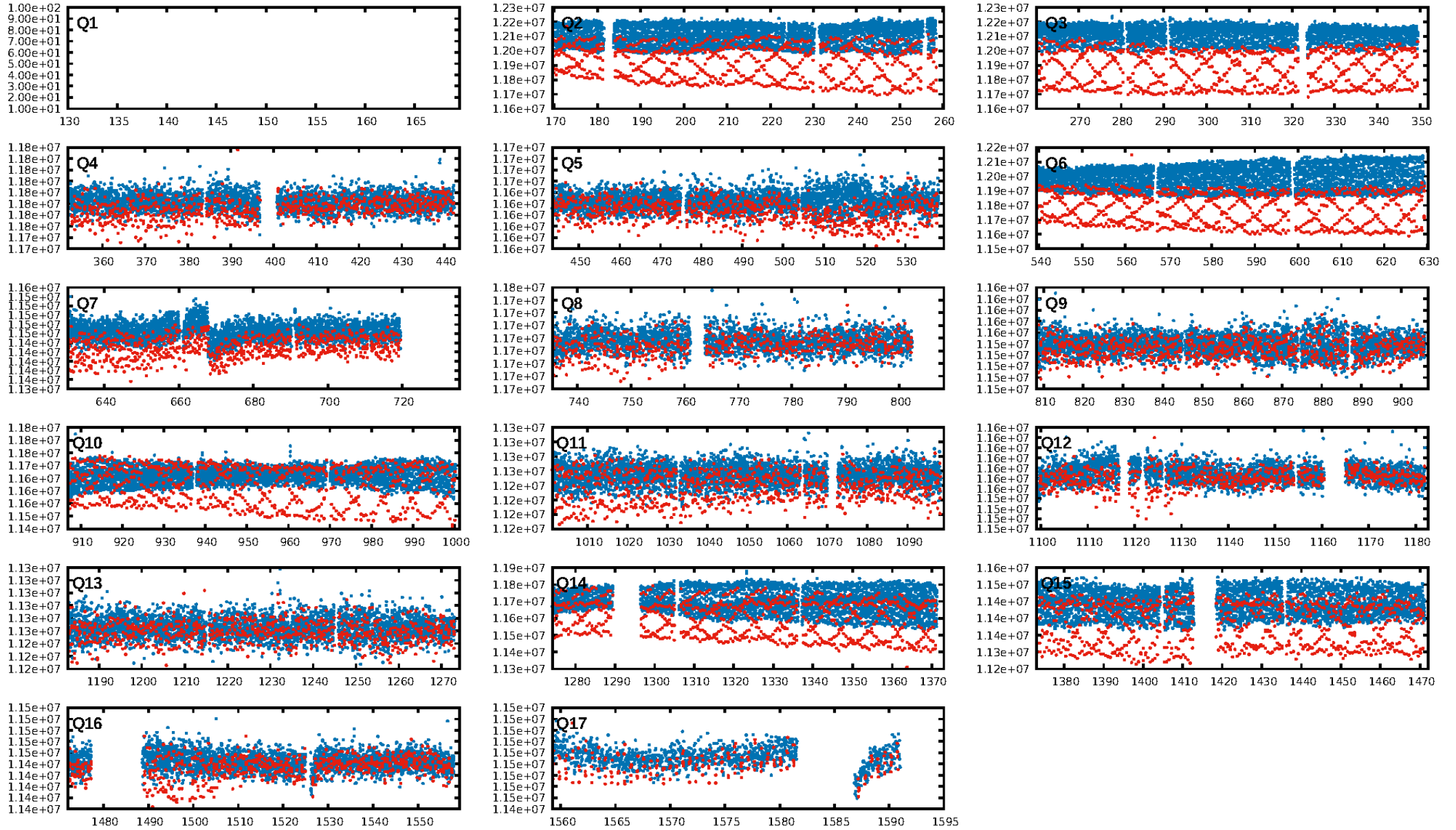
DV Fit Results:

Period = 0.59428 [0.00000] d
Epoch = 131.8379 [0.0004] BKJD
Rp/R* = 0.0417 [0.1095]
a/R* = 1.52 [0.52]
b = 1.00 [0.17]
Seff = 8346.19 [3223.29]
Teq = 2437 [235] K
Rp = 5.12 [13.52] Re
a = 0.0147 [0.0036] AU
Ag = 0.06 [0.35] [-2.66σ]
Teffp = 1896 [2594] K [-0.21σ]

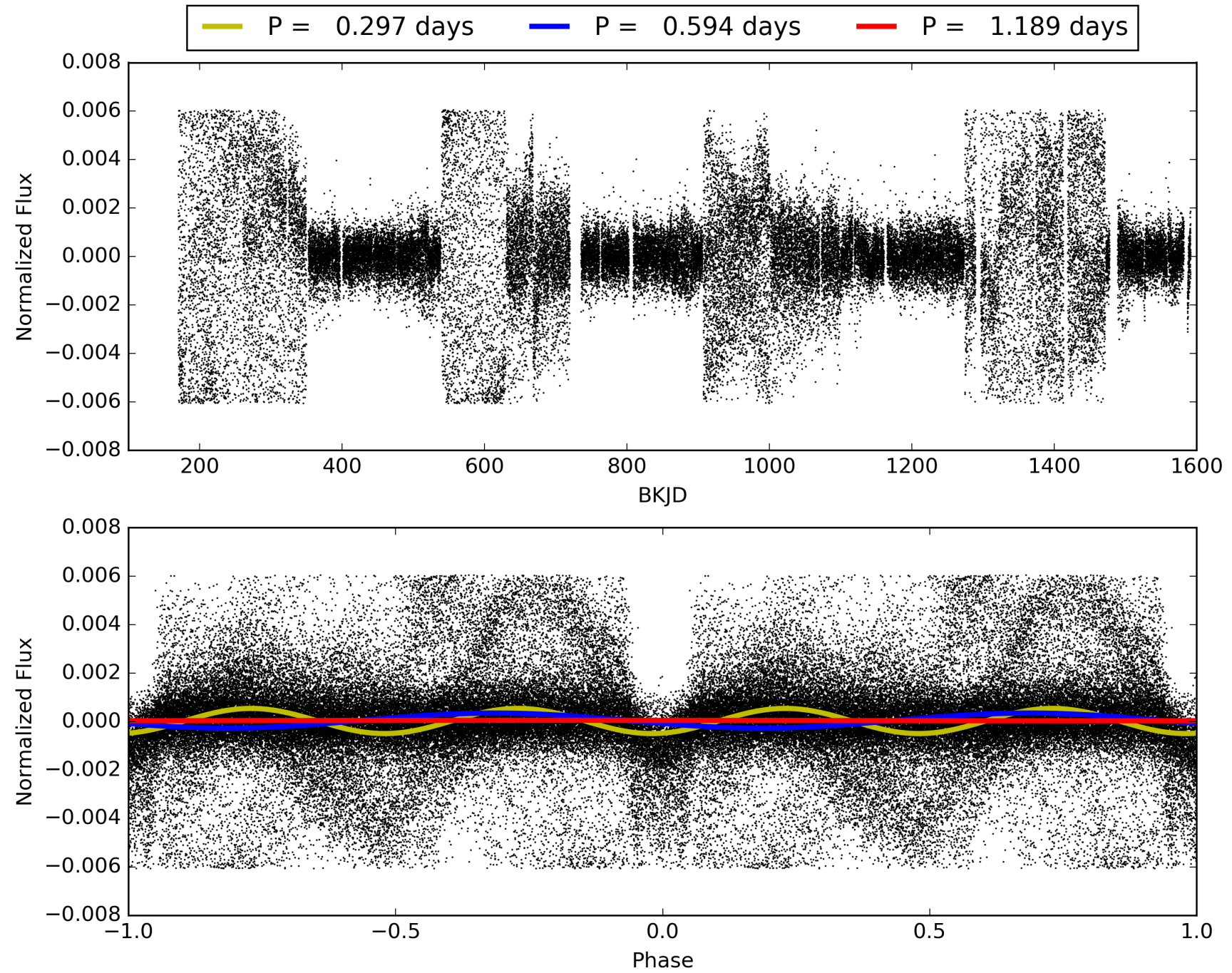
DV Diagnostic Results:

ShortPeriod-sig: N/A
LongPeriod-sig: N/A
ModelChiSquare2-sig: N/A
ModelChiSquareGof-sig: N/A
Bootstrap-pfa: 0.00e+00
RollingBand-fgt: 0.99 [2126/2151]
GhostDiagnostic-chr: -0.337
Centroid-sig: N/A
Centroid-so: 21.182 arcsec [61.52σ]
OotOffset-rm: 5.882 arcsec [10.33σ]
KicOffset-rm: 8.506 arcsec [121.24σ]
OotOffset-st: 4/4/4/0 [12]
KicOffset-st: 4/4/4/0 [12]
DiffImageQuality-fgm: 1.00 [12/12]
DiffImageOverlap-fno: 1.00 [16/16]

TCE 003954211-01, PDC Light Curves

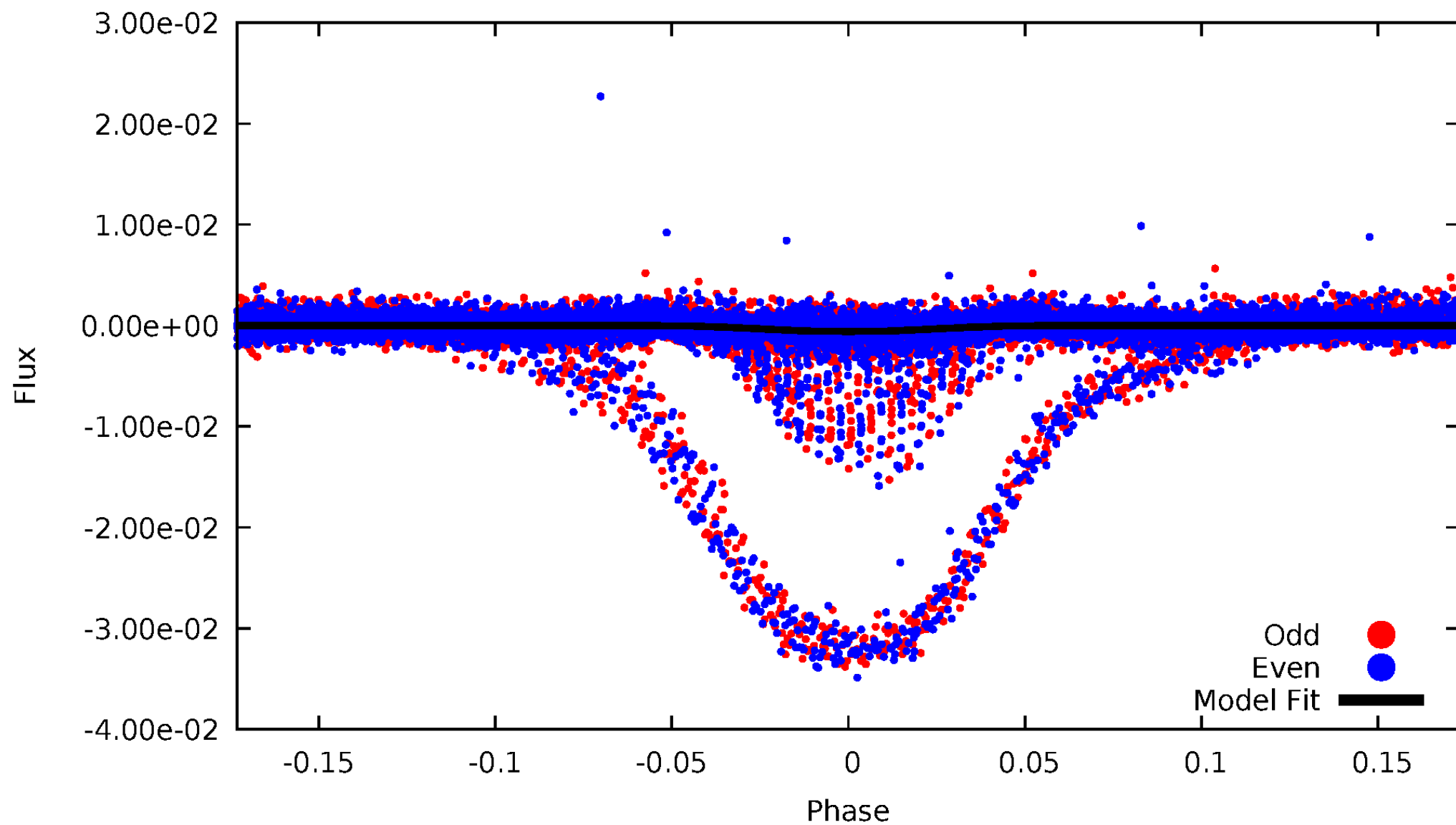


TCE 003954211-01



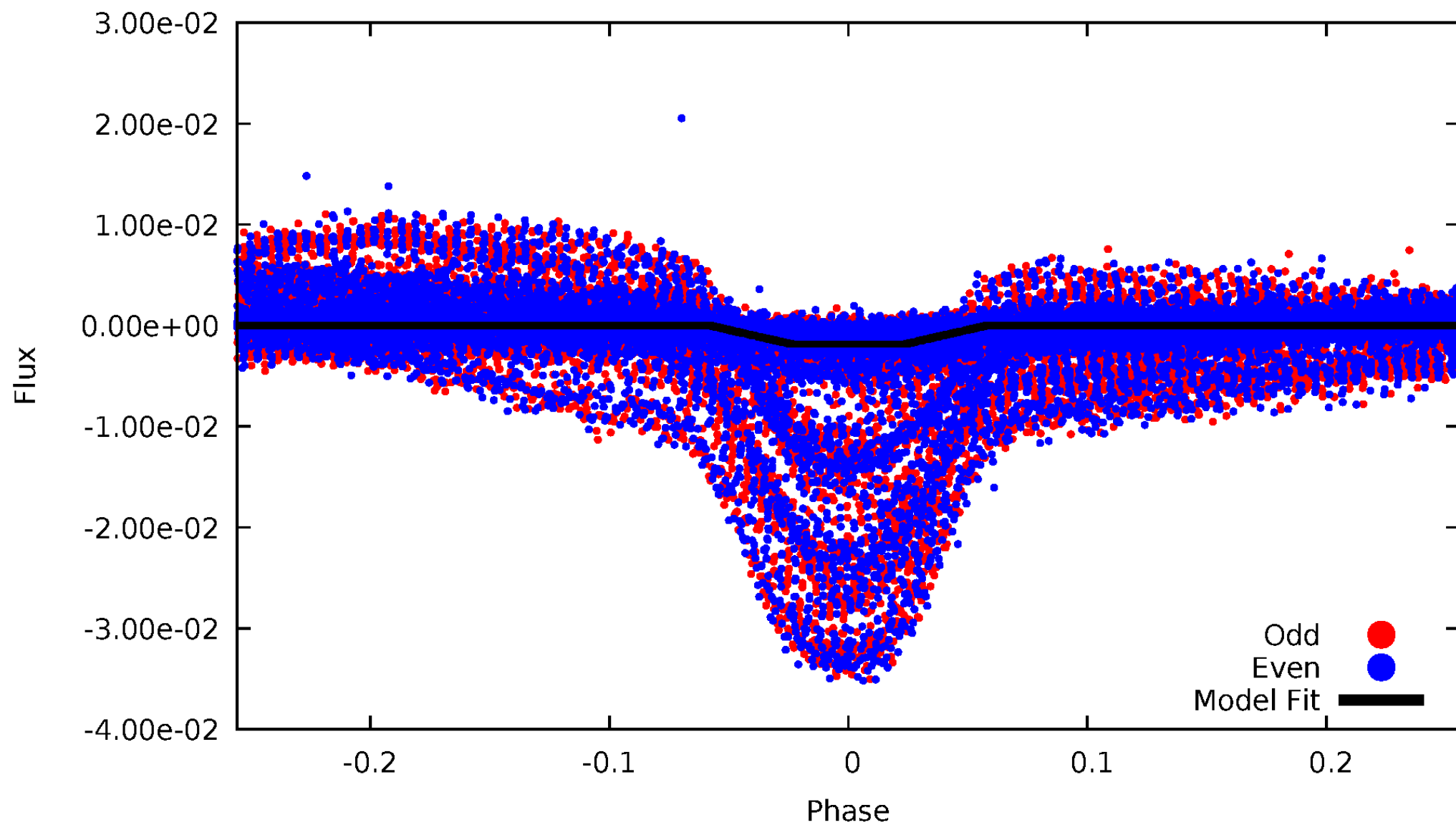
DV Odd/Even

TCE 003954211-01



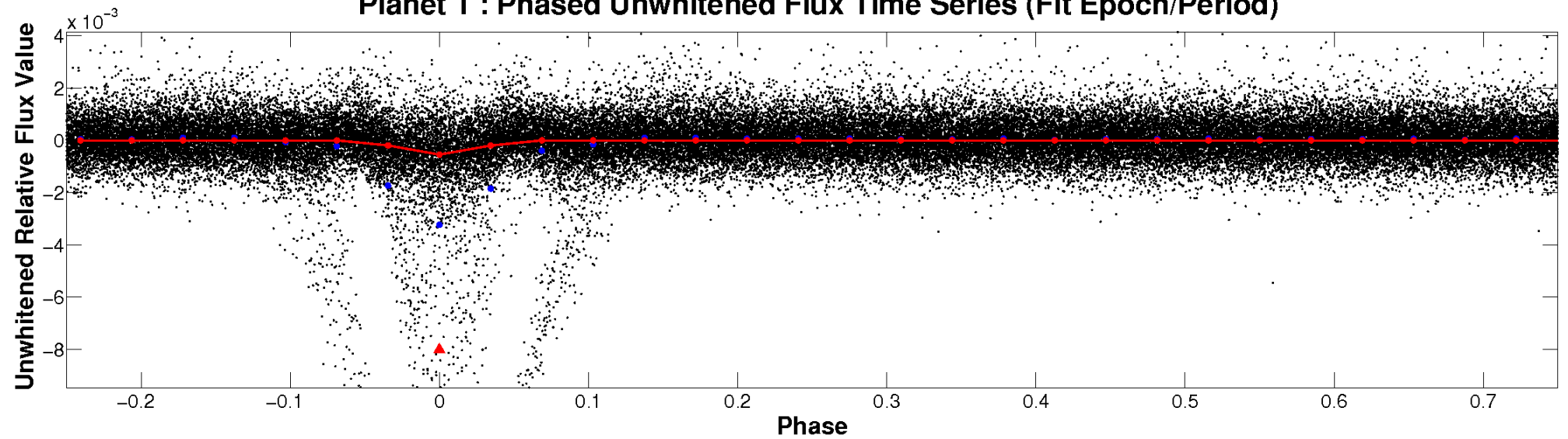
ALT Odd/Even

TCE 003954211-01

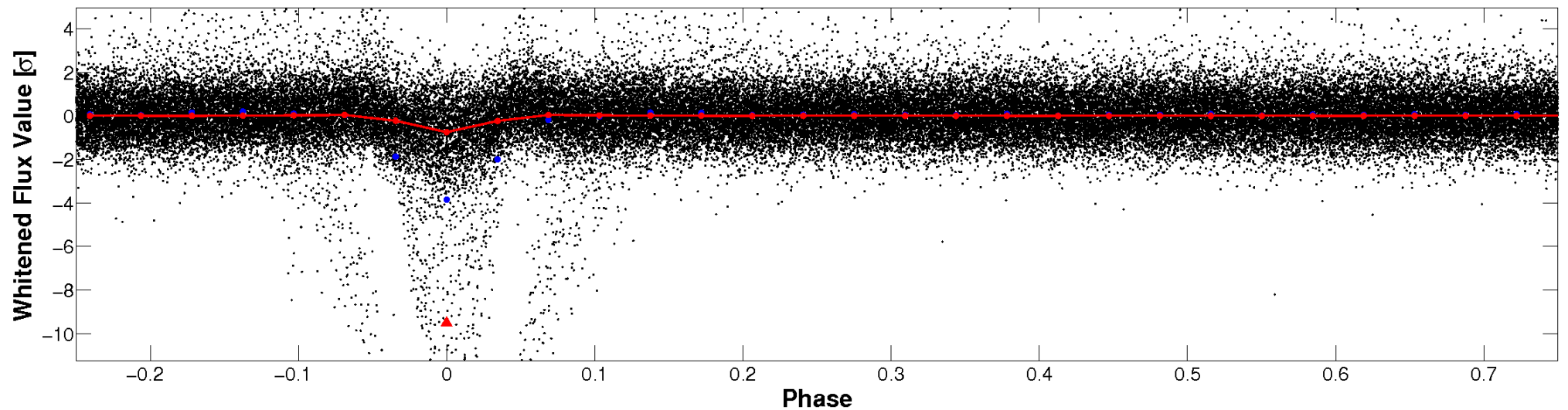


Non-Whitened Vs. Whitened Light Curve

Planet 1 : Phased Unwhitened Flux Time Series (Fit Epoch/Period)

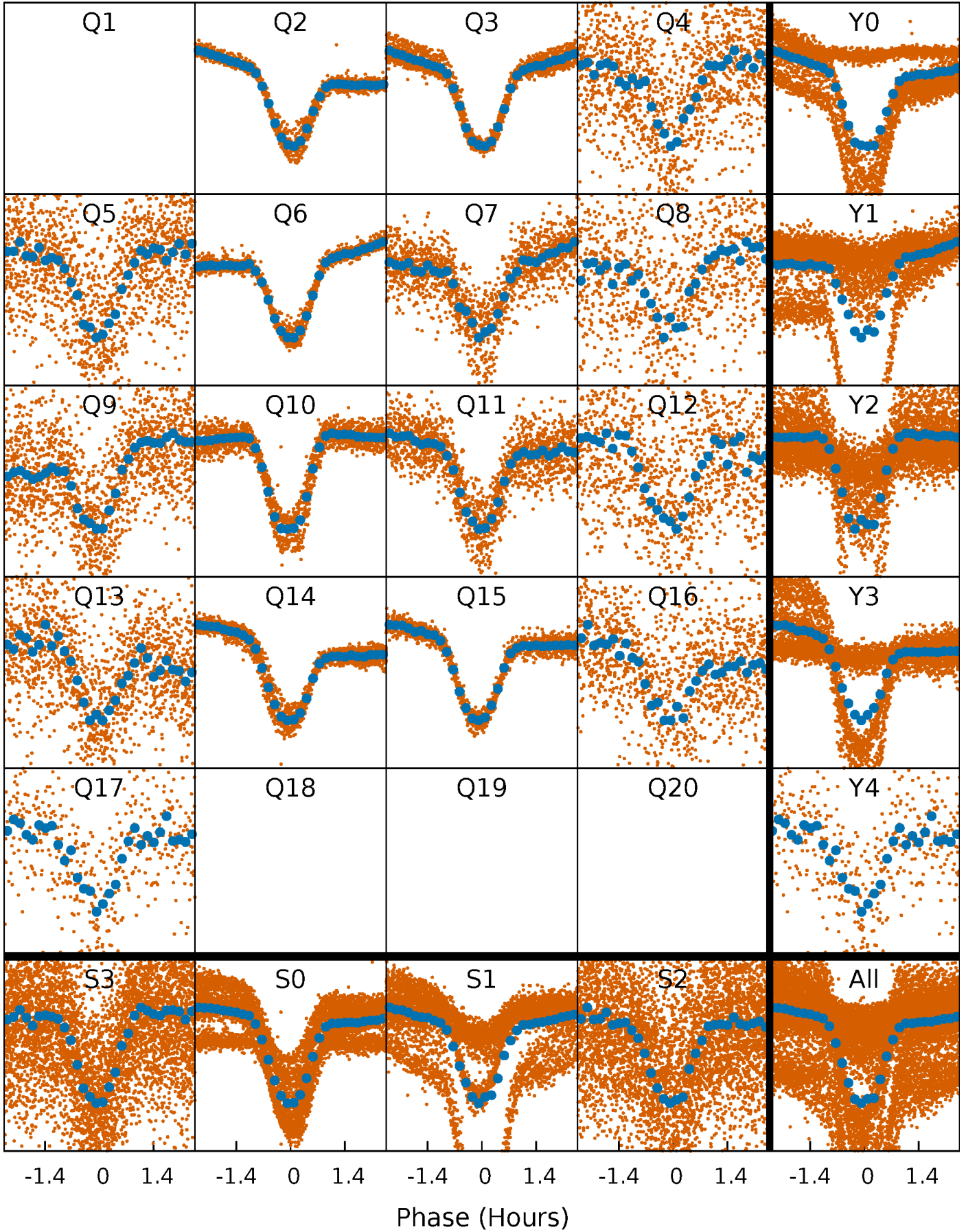


Planet 1 : Phased Whitened Flux Time Series (Fit Epoch/Period)



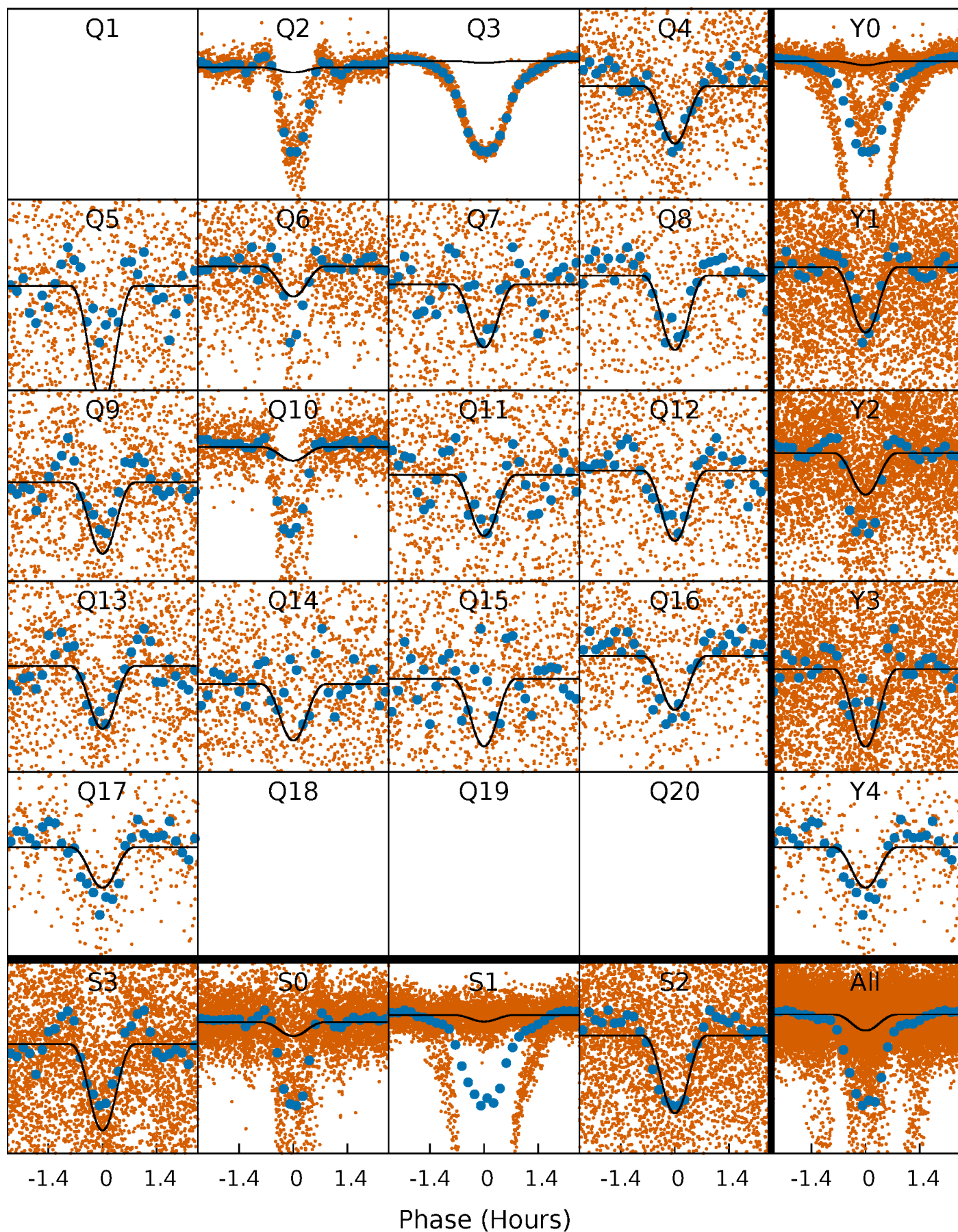
PDC Quarter-Phased Transit Curves

TCE 003954211-01 P= 0.594276 Days $T_0=131.837901$ (BKJD)



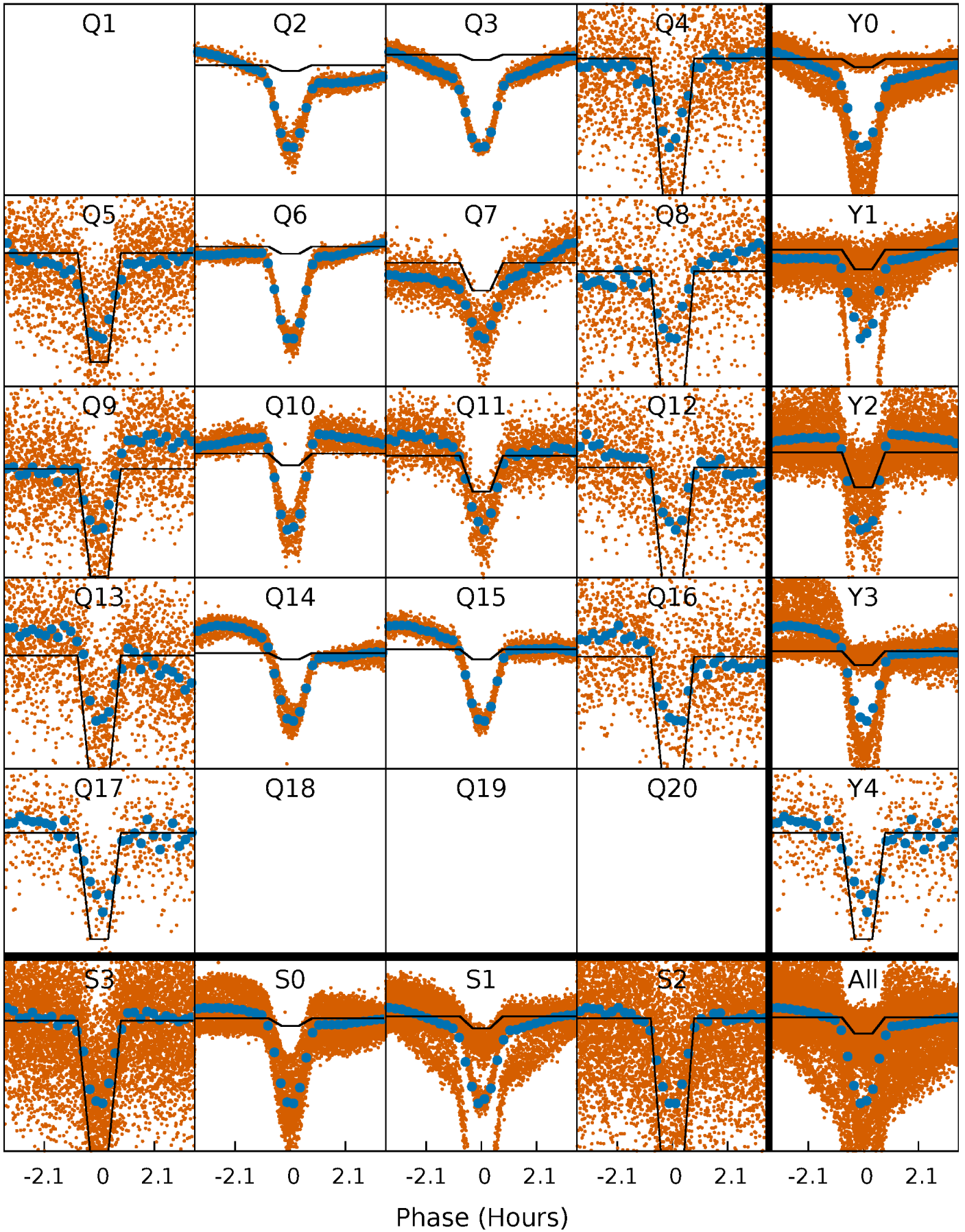
DV Quarter-Phased Transit Curves

TCE 003954211-01 P= 0.594276 Days $T_0=131.837901$ (BKJD)



Alt. Detrend Quarter-Phased Transit Curves

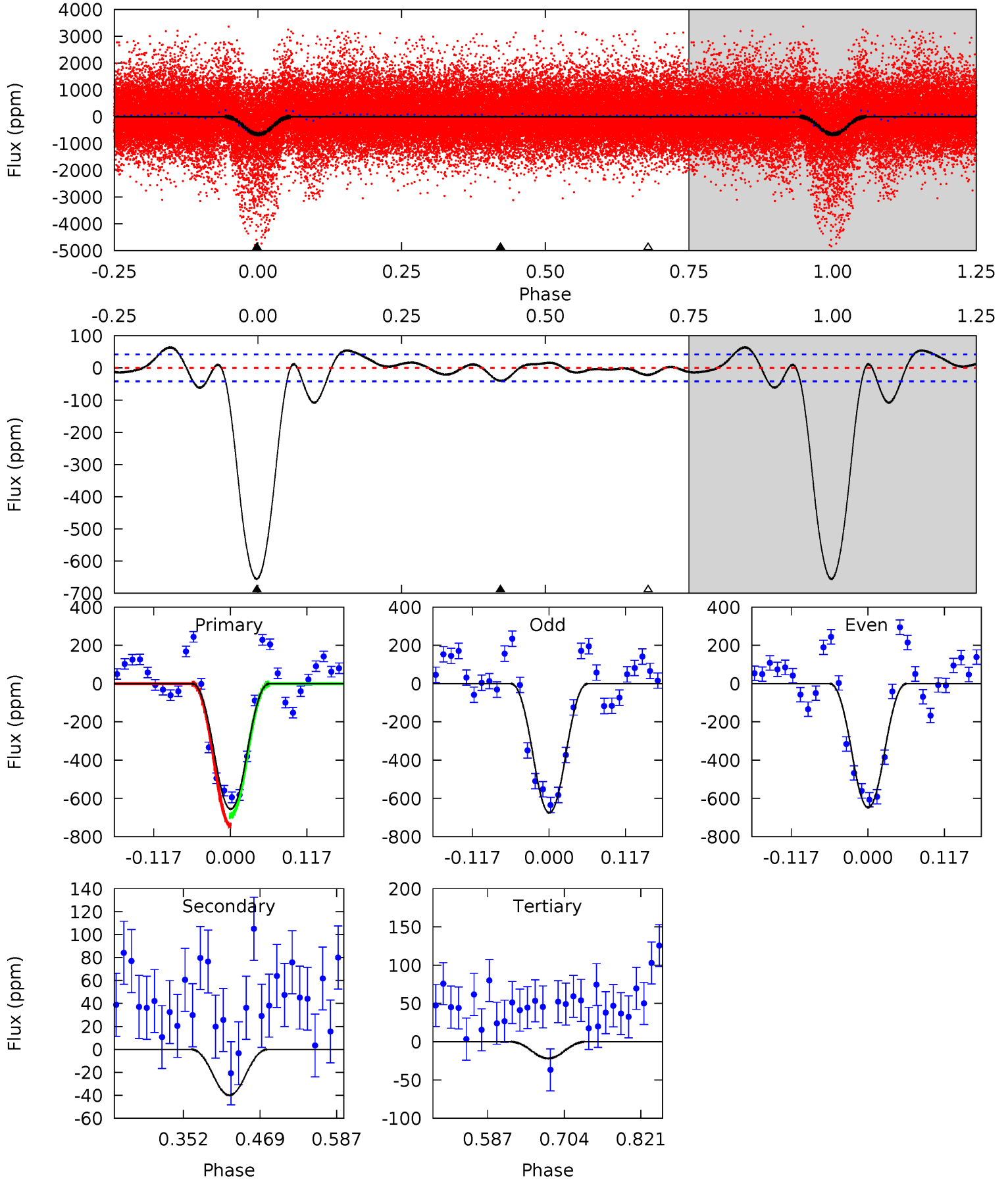
TCE 003954211-01 P= 0.594273 Days $T_0=131.839992$ (BKJD)



DV Model-Shift Uniqueness Test

003954211-01, P = 0.594276 Days, E = 131.837901 Days

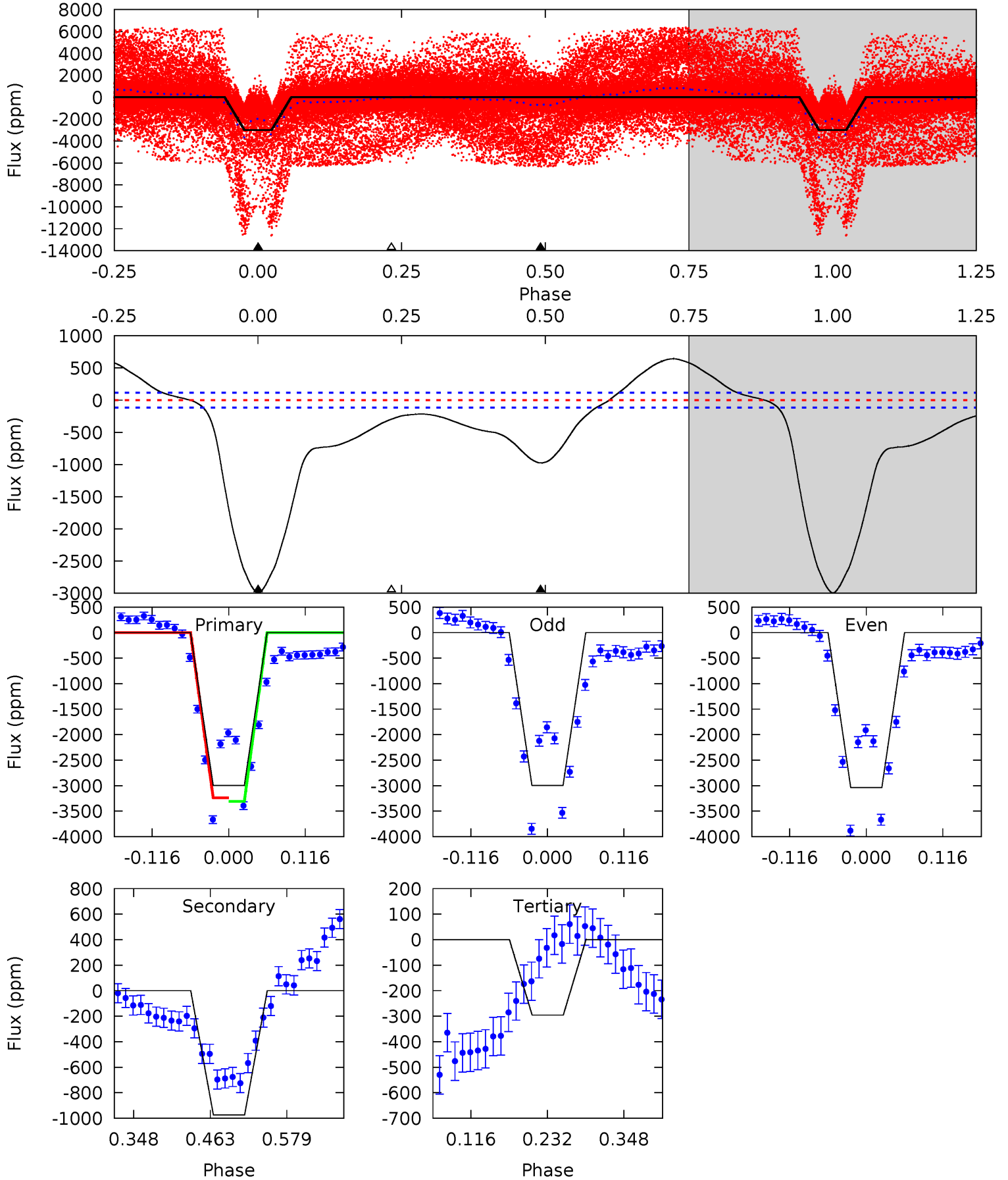
| Pri | Sec | Ter | Pos | FA ₁ | FA ₂ | F _{Red} | Pri-Ter | Pri-Pos | Sec-Ter | Sec-Pos | Odd-Evn | DMM | Shape | TAT |
|------|------|------|-----|-----------------|-----------------|------------------|---------|---------|---------|---------|---------|------|-------|------|
| 71.1 | 4.32 | 2.36 | 0 | 4.53 | 1.57 | 2.48 | 68.7 | 71.1 | 1.96 | 4.32 | 1.49 | 5.99 | 0.09 | 2.83 |



Alt Model-Shift Uniqueness Test

003954211-01, P = 0.594273 Days, E = 131.839992 Days

| Pri | Sec | Ter | Pos | FA ₁ | FA ₂ | F _{Red} | Pri-Ter | Pri-Pos | Sec-Ter | Sec-Pos | Odd-Evn | DMM | Shape | TAT |
|-------|------|------|-----|-----------------|-----------------|------------------|---------|---------|---------|---------|---------|------|-------|-----|
| 117.1 | 38.1 | 11.6 | 0 | 4.53 | 1.57 | 16.2 | 105.5 | 117.1 | 26.6 | 38.1 | 0.86 | 2.67 | 0.18 | 0 |



Stellar Parameters For KIC 003954211

| | $T_{\text{eff}}(K)$ | $\log(g)$ | [Fe/H] | $R (R_{\odot})$ | $M(M_{\odot})$ | $p_{\star} (\text{g}\cdot\text{cm}^{-3})$ |
|--------|----------------------|---------------------------|---------------------------|---------------------------|---------------------------|---|
| | 6321^{+174}_{-239} | $4.417^{+0.060}_{-0.192}$ | $0.070^{+0.250}_{-0.300}$ | $1.125^{+0.335}_{-0.112}$ | $1.207^{+0.141}_{-0.173}$ | $1.195^{+0.361}_{-0.573}$ |
| | +3%/-4% | +1%/-4% | +357%/-429% | +30%/-10% | +12%/-14% | +30%/-48% |
| Source | PHO1 | KIC0 | KIC0 | DSEP | | |

KIC = Kepler Input Catalog; PHO = Photometry; SPE = Spectroscopy; AST = Asteroseismology
 TRA = Transits; DESP = Dartmouth Models; MULT = Multiple Models

Secondary Eclipse Parameters for KIC 003954211-01 / KOI 8090.01

| Detrend | Depth (ppm) | $R_p (R_{\oplus})$ | $T_{\text{max}} (K)$ | $T_{\text{obs}} (K)$ | A_{obs} |
|---------|---------------|--------------------------|----------------------|------------------------|---------------------------|
| DV | -40 ± 9 | $11.32^{+11.25}_{-7.74}$ | 3455^{+225}_{-168} | -3257^{+5919}_{-170} | $0.036^{+0.324}_{-0.027}$ |
| Alt. | -975 ± 26 | $12.73^{+11.61}_{-8.51}$ | 3469^{+229}_{-188} | 3522^{+2535}_{-6456} | $0.699^{+5.803}_{-0.505}$ |

T_{max} = Theoretical Maximum Planetary Temperature

T_{obs} = Observed Planetary Temperature (Assuming $A=0.3$)

A_{obs} = Observed Albedo (Assuming $T=0$)

If a secondary eclipse is present, the system is likely an EB if $T_{\text{obs}} \gg T_{\text{max}}$ AND $A_{\text{obs}} \gg 1.0$

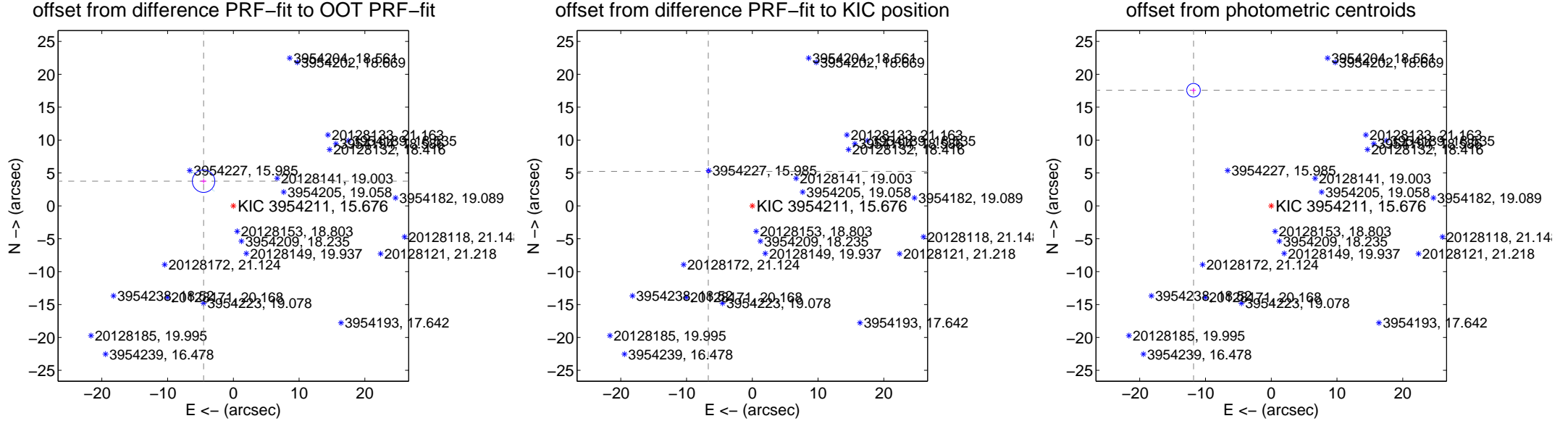
DV Centroid Data

Supplemental centroid analysis for 003954211-01. Kepler magnitude: 15.68. Transit SNR 36.91

There are 12 quarters with good PRF difference image offsets

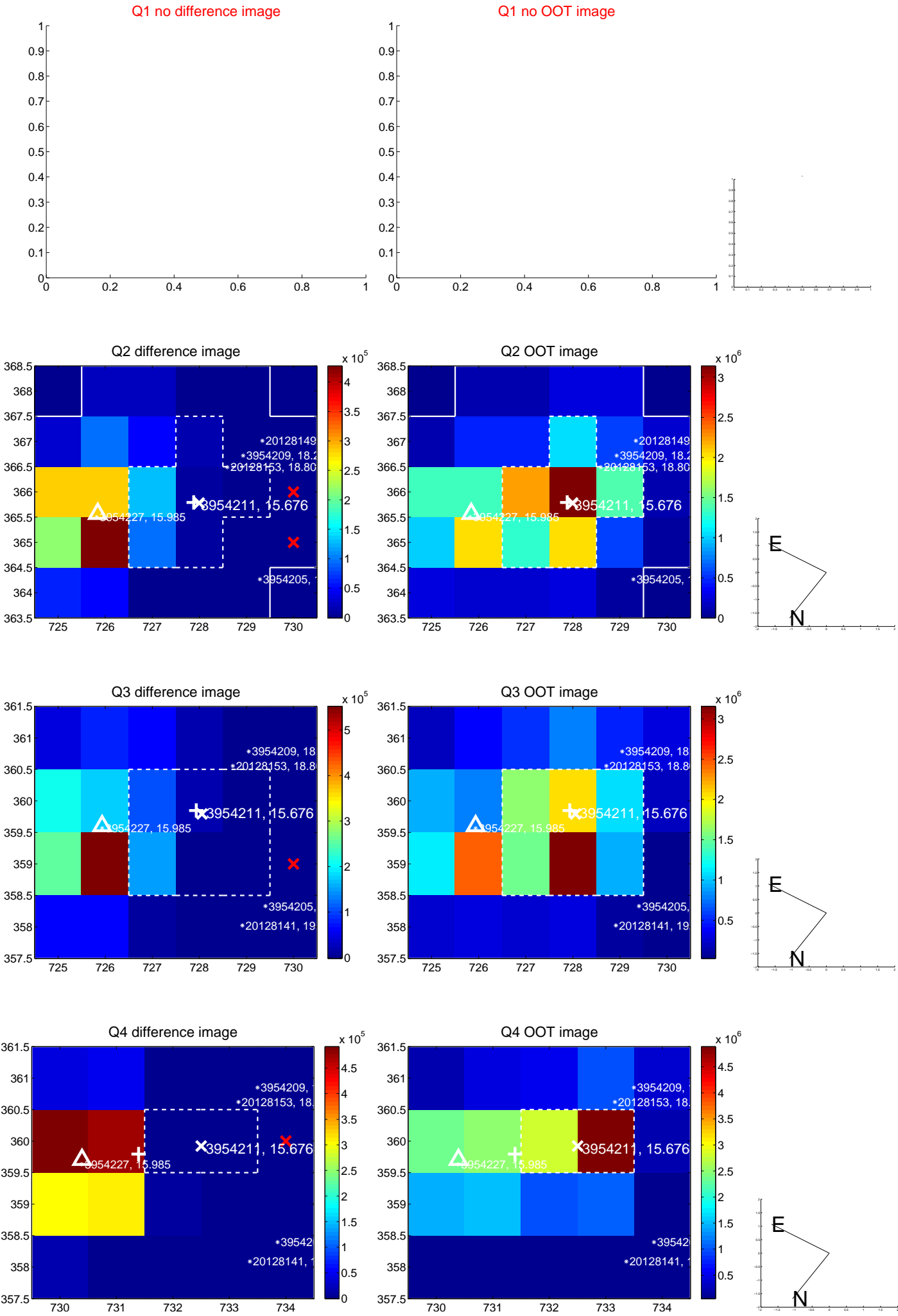
The OOT PRF centroid is offset from the target star catalog position by about 4.60 arcsec so the offset from difference PRF-fit to OOT-fit may be invalid.

| | Distance in arcsec | Distance / σ | Δ RA | Δ Dec |
|---|--------------------|---------------------|-------------------|-------------------|
| PRF-fit source offset from OOT | 5.882 ± 0.570 | 10.33 | 4.528 ± 0.436 | 3.755 ± 0.373 |
| PRF-fit source offset from KIC position | 8.506 ± 0.070 | 121.24 | 6.695 ± 0.071 | 5.247 ± 0.068 |
| photometric centroid source offset | 21.18 ± 0.34 | 61.53 | 11.82 ± 0.36 | 17.58 ± 0.34 |

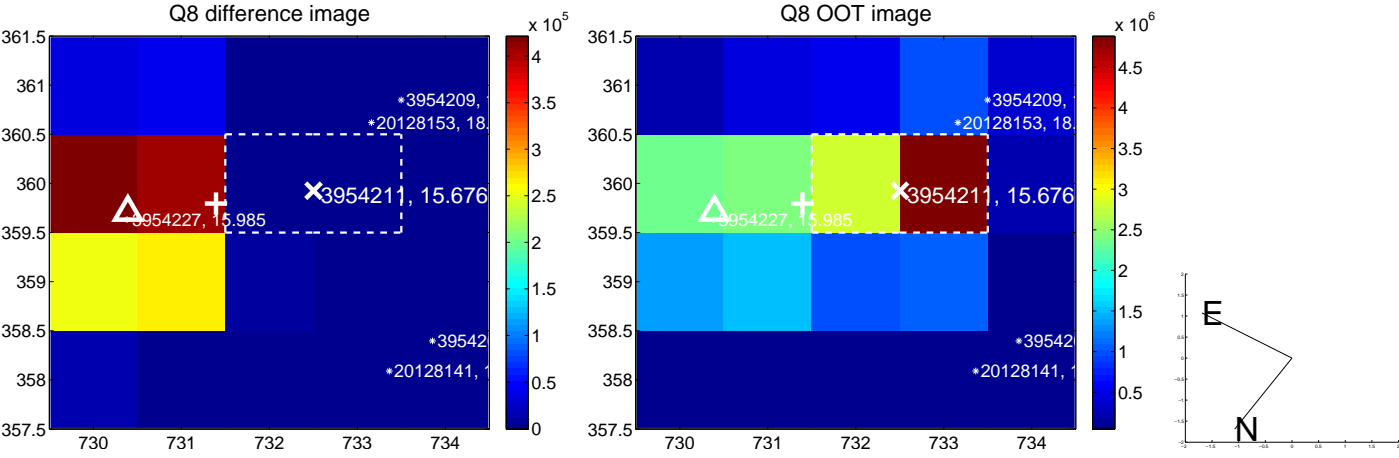
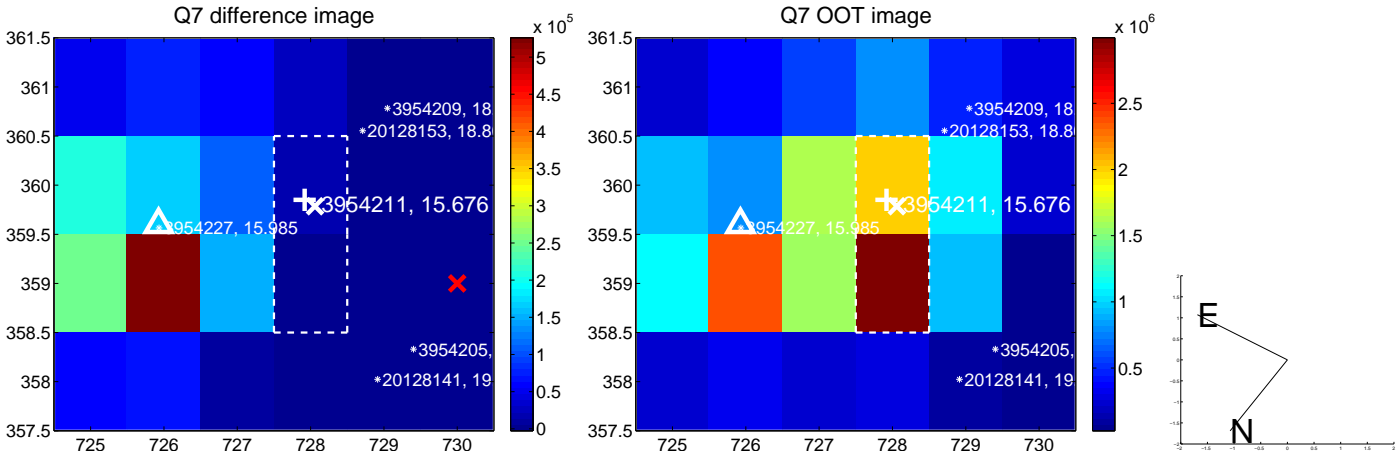
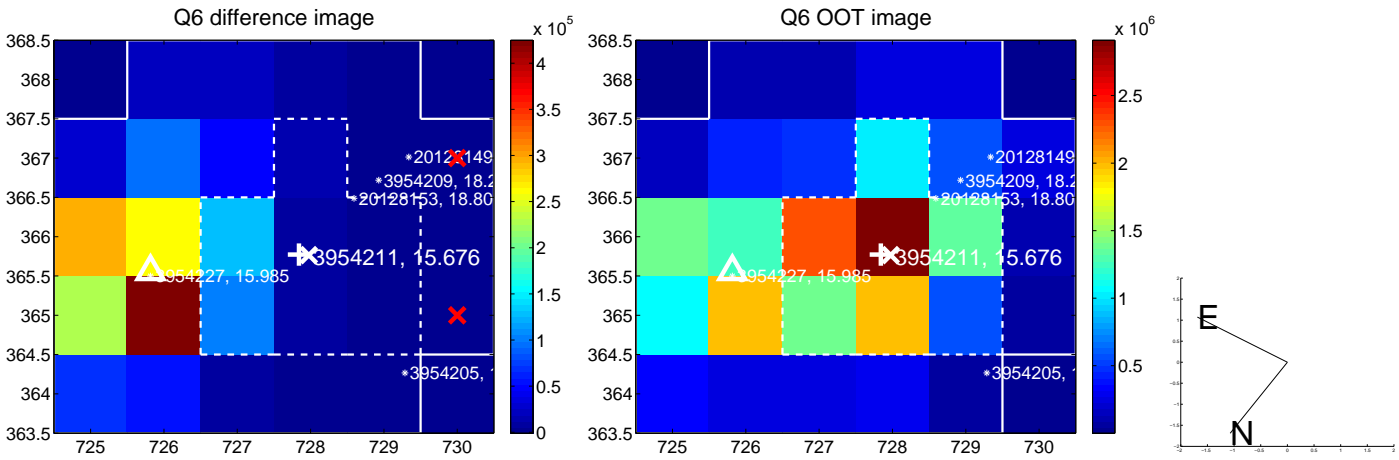
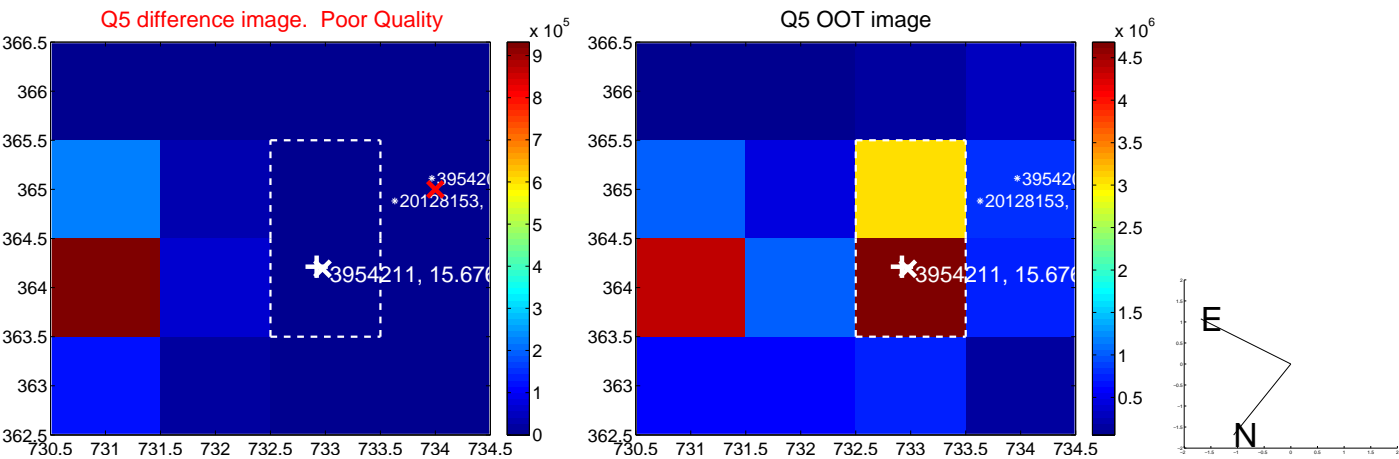


Centroid source offsets from the target star reconstructed from PRF and photometric centroids. Sky blue crosses: good quarterly centroid offsets; Vermillion crosses: bad quarterly centroid offsets; magenta cross: average over quarters. Length of the crosses: one- σ uncertainty. Blue circle: three- σ . Red *: target star. Blue *: Other stars. Text next to a star gives its KIC ID and kepmag. KIC IDs > 15,000,000 are from the UKIRT catalog.

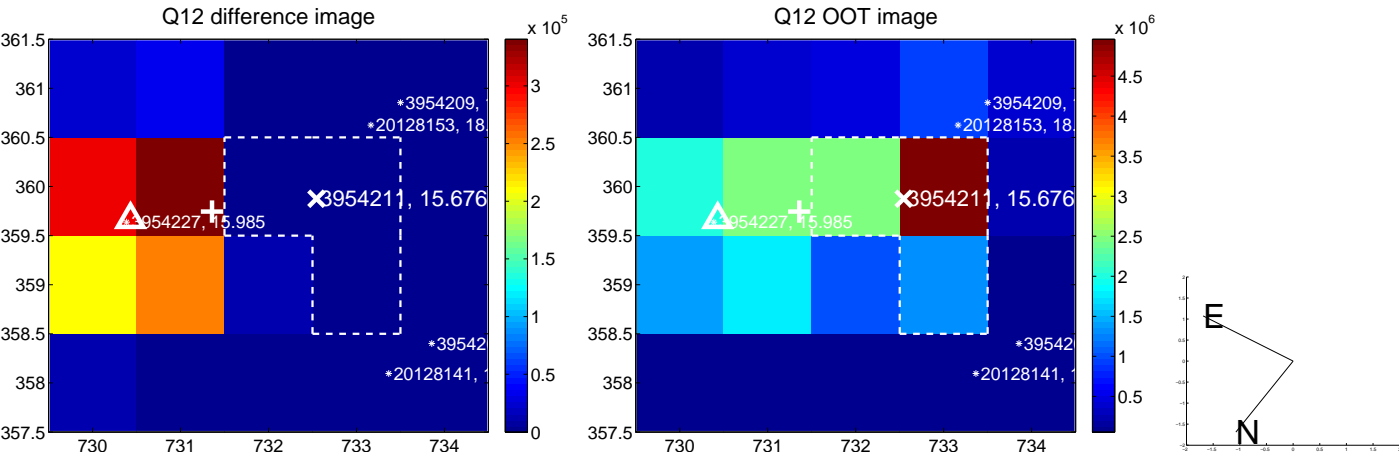
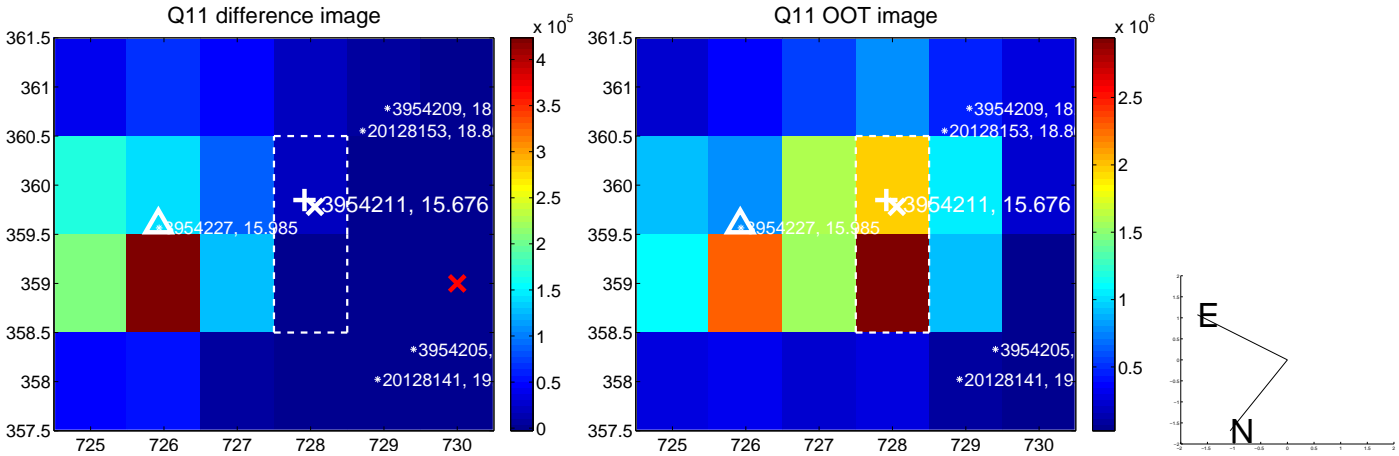
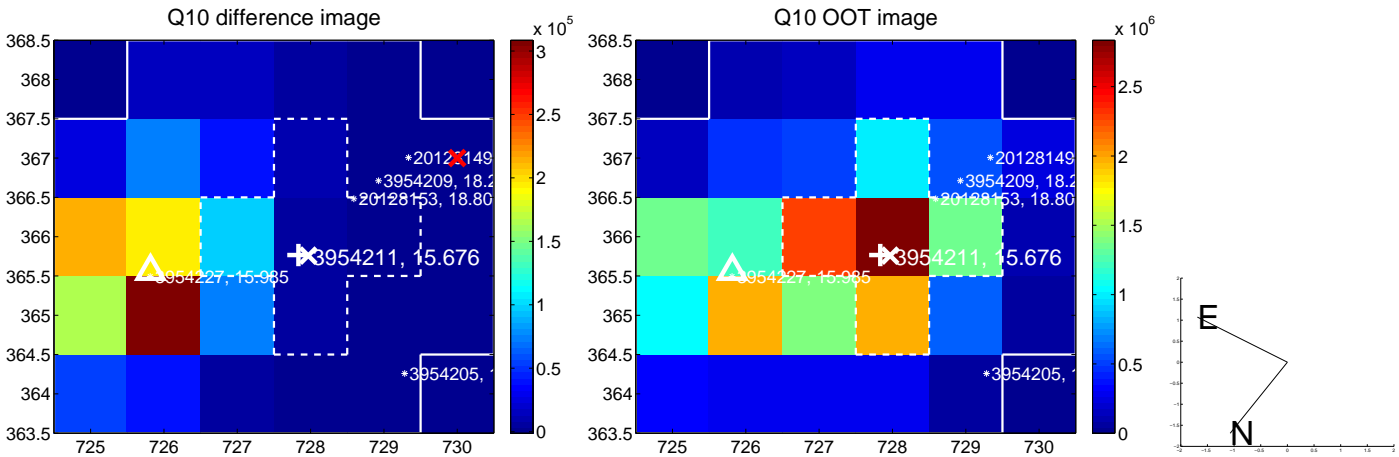
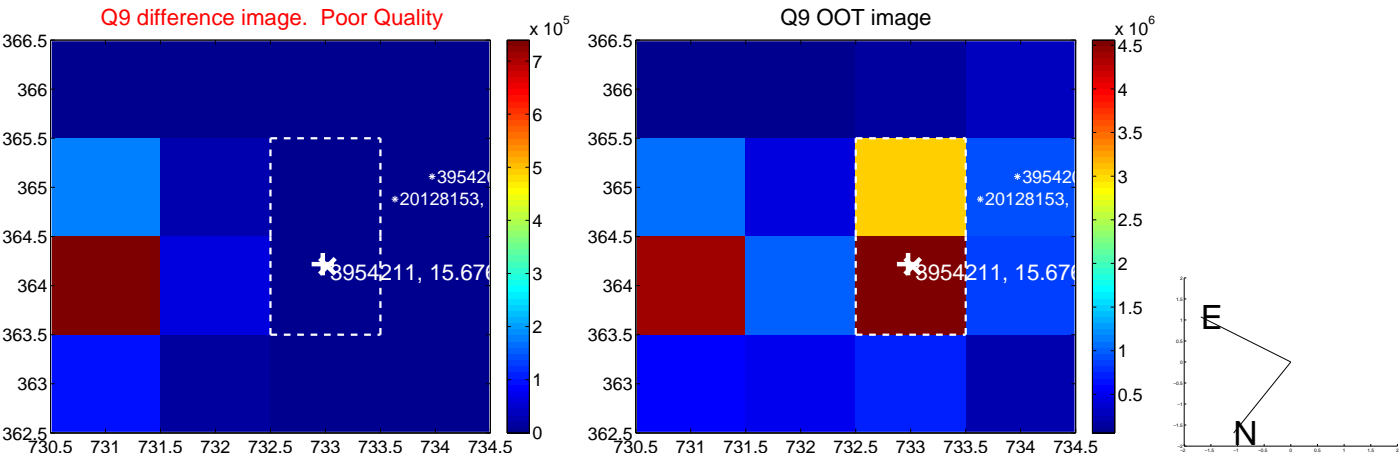
white \times : KIC target position; $+$: OOT centroid; \triangle : difference centroid. red \times : large negative pixel value.



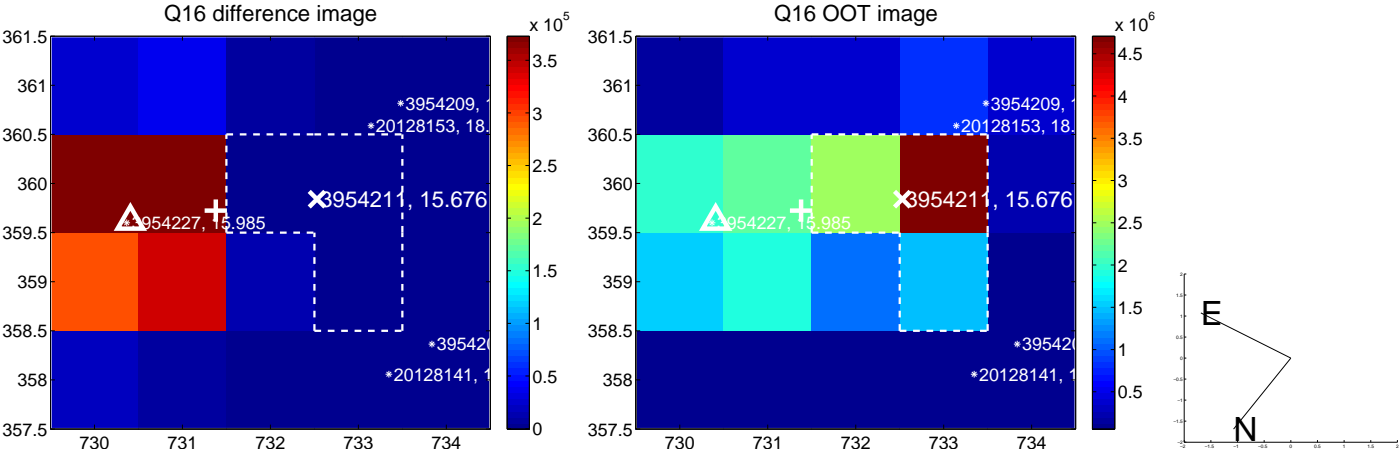
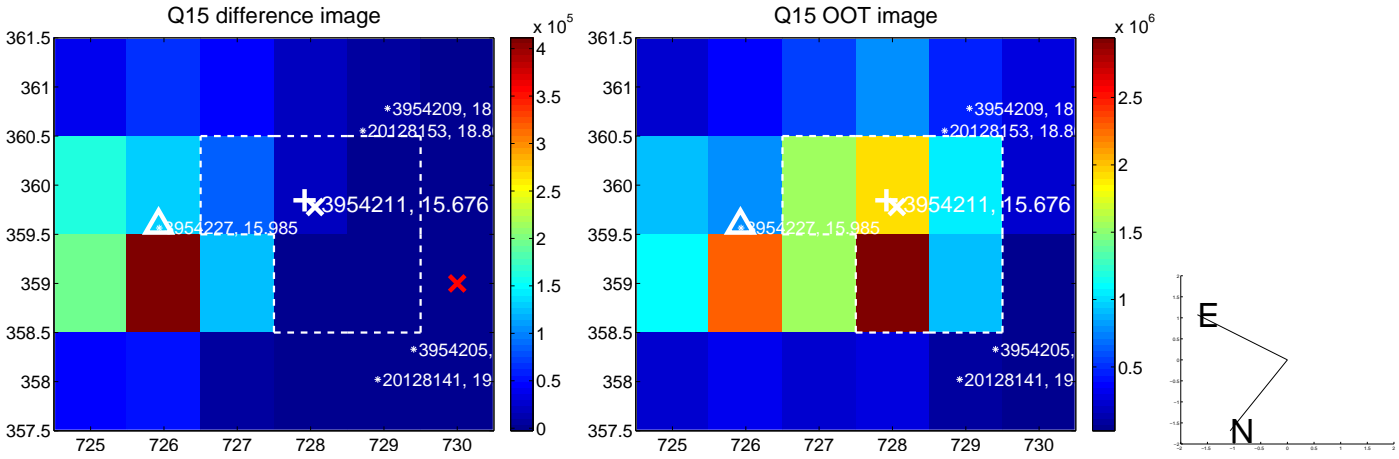
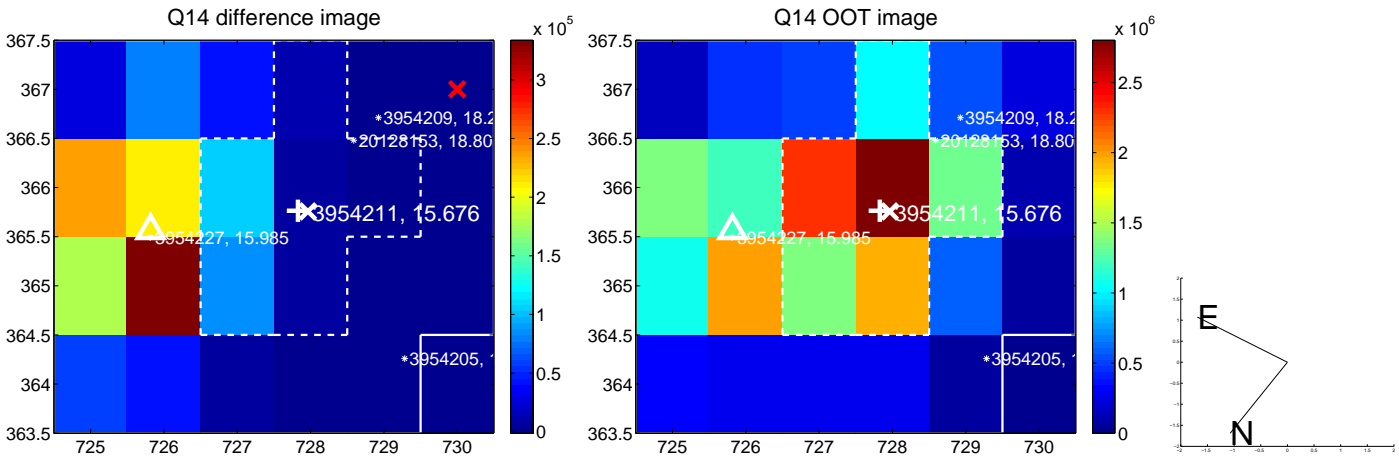
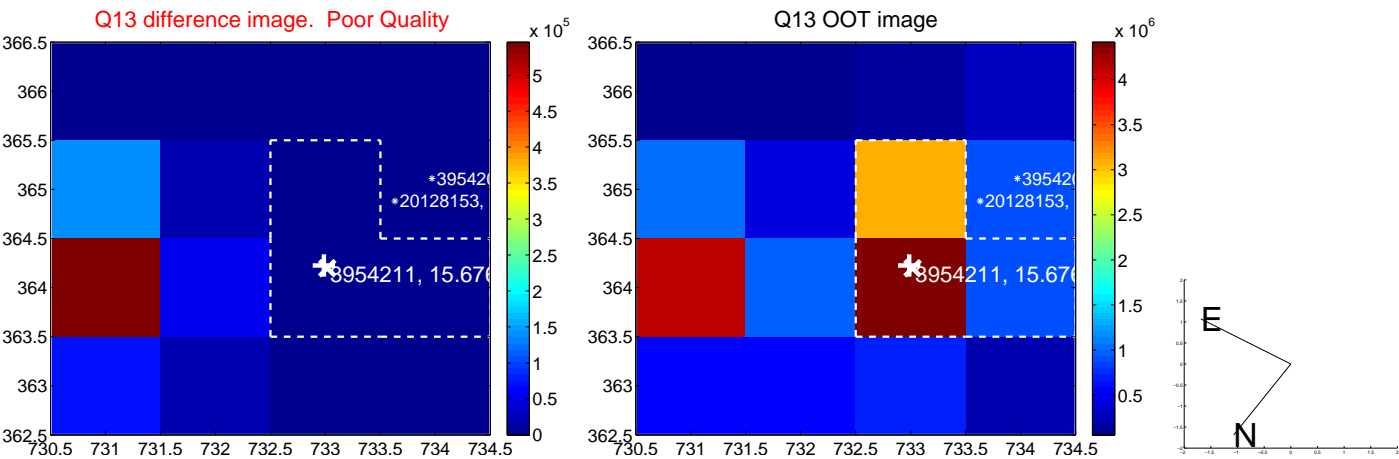
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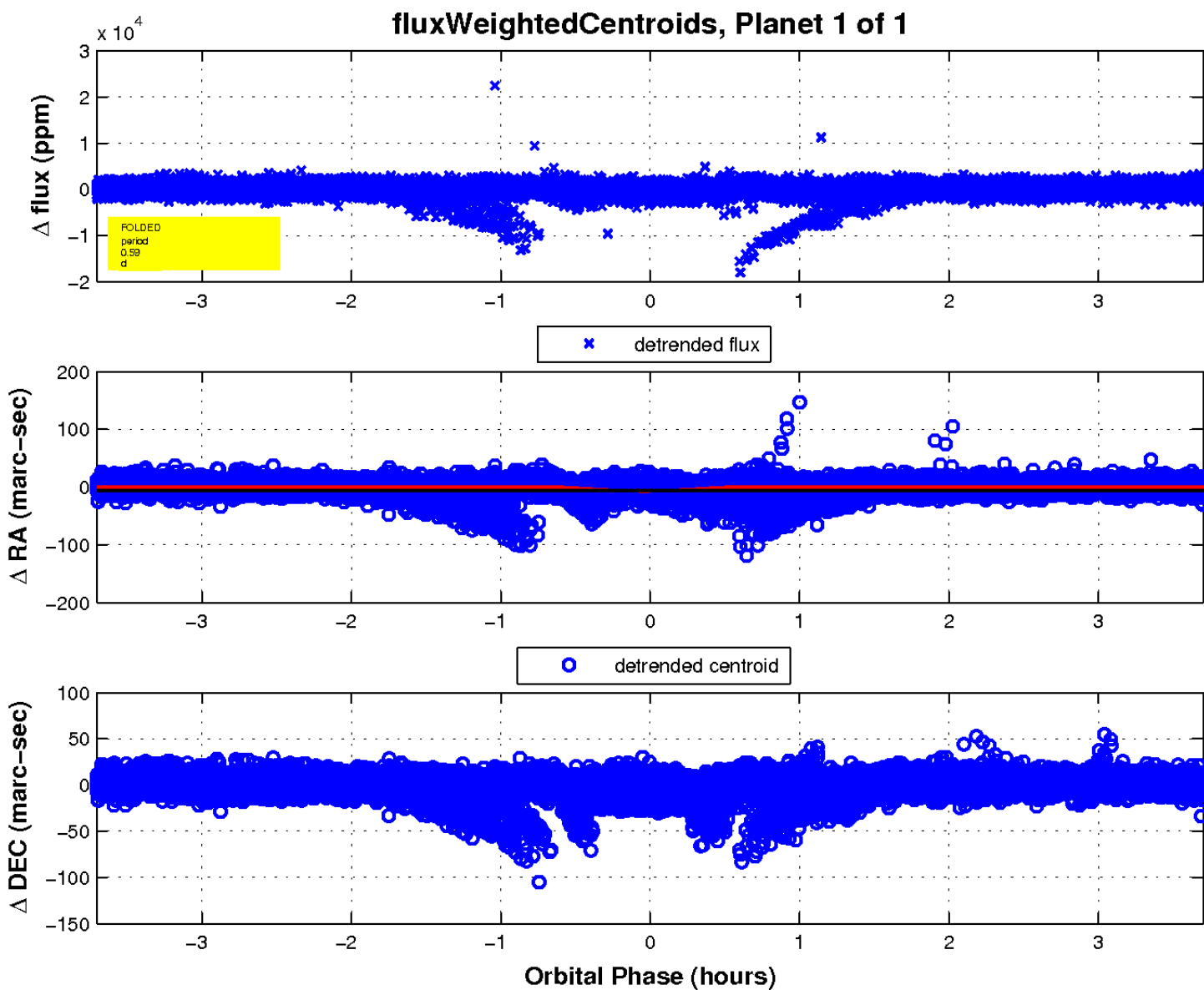
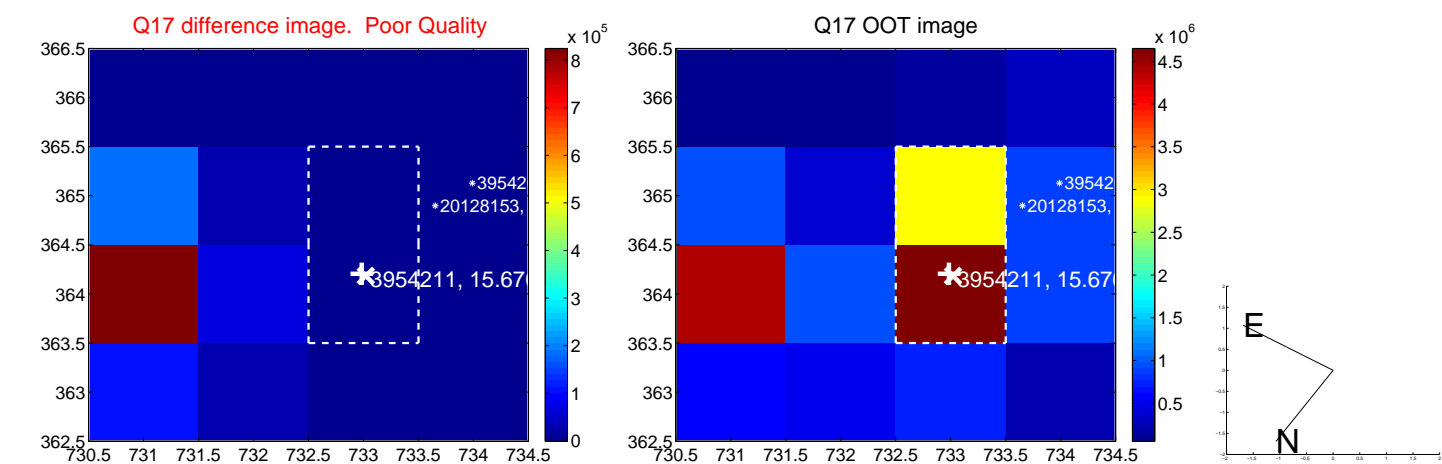
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UKIRT Image

Declination

